

# Helping to Improve the Analysis of SNO+ Data with Machine Learning

**Mark Anderson**

*on behalf of the SNO+ collaboration*

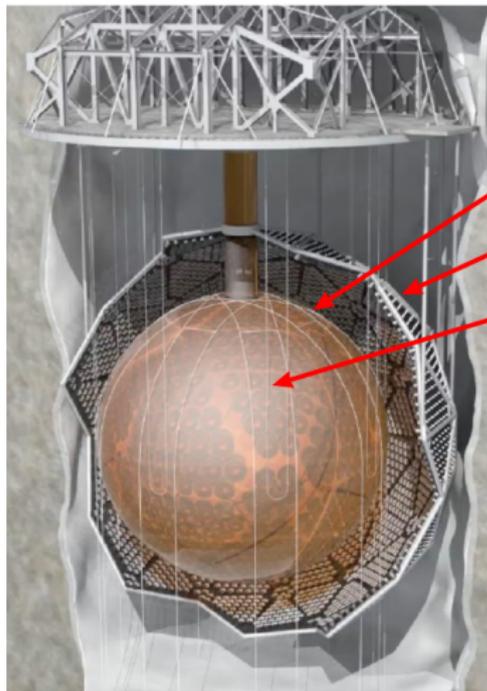


WNPPC · February 17<sup>th</sup>, 2019

# Overview

- SNO+ detector and physics goals
- Machine learning basics
- Applications of machine learning to SNO+
  - Antineutrinos
  - Reconstruction

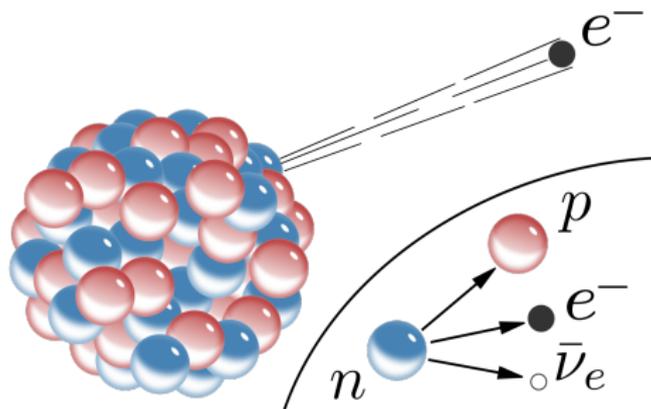
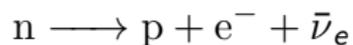
# The SNO+ Detector



- Multipurpose neutrino experiment
- 2km underground
- 12m diameter acrylic vessel
- ~10,000 PMTs
- 900Mg water / 780Mg scintillator
- Physics goals include
  - Neutrinoless double beta decay
  - Low energy solar neutrinos
  - **Antineutrinos**

# SNO+ Antineutrinos

- Antineutrinos originate from several sources
  - $\beta$ -decay in nuclear reactors (reactor neutrinos)
  - $\beta$ -decay of radioactive isotopes in the Earth (geo-neutrinos)

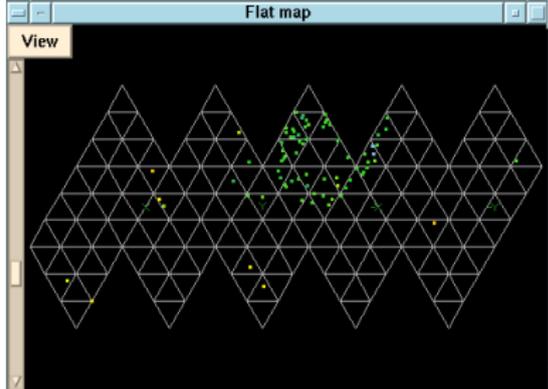
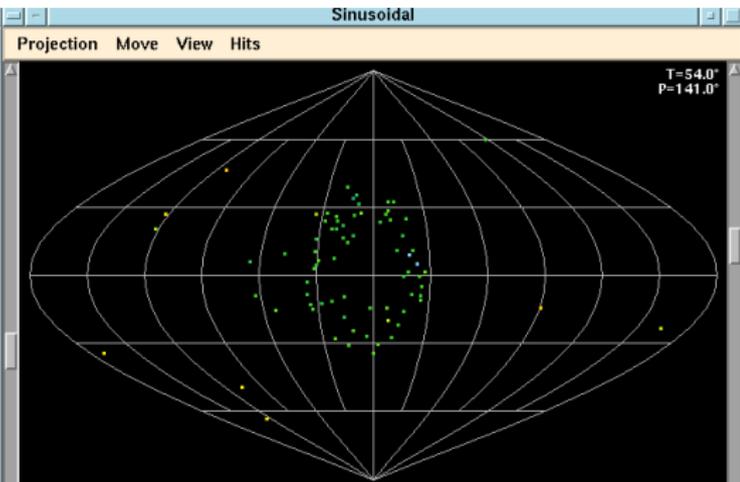
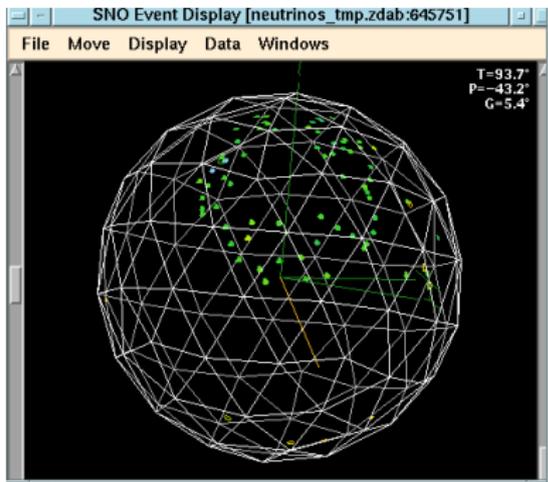


# SNO+ Antineutrinos

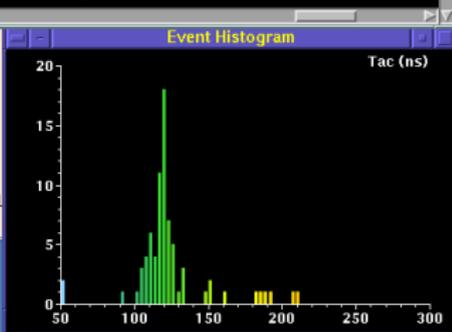
- Goal is to detect antineutrino signal in water
- Detected via inverse beta decay



- Coincident signal
  - Positron deposits energy, annihilates (*prompt* event)
  - Neutron is captured later by a proton (*delayed* event)
    - Emits 2.2MeV  $\gamma$
  - Signals present as rings on the detector
- Antineutrino signals are difficult to detect in water
  - Low energy deposited
  - Dominated by other backgrounds in the detector

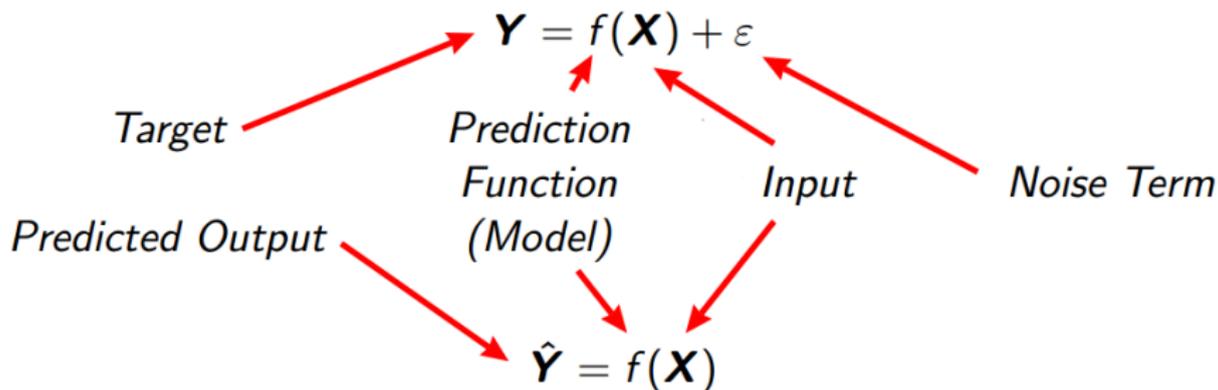


GTID: 645751  
 Evt Num: 646705  
 Run Num: 5463  
 Date: 10/09/1999  
 Time: 04:42:26.0497823  
 Prev/Next: 141 ms / 789 ms  
 Trigger: 20LB,20,100H,100M,100L  
 Pk/Int/Dif: 12 / 120 / 0

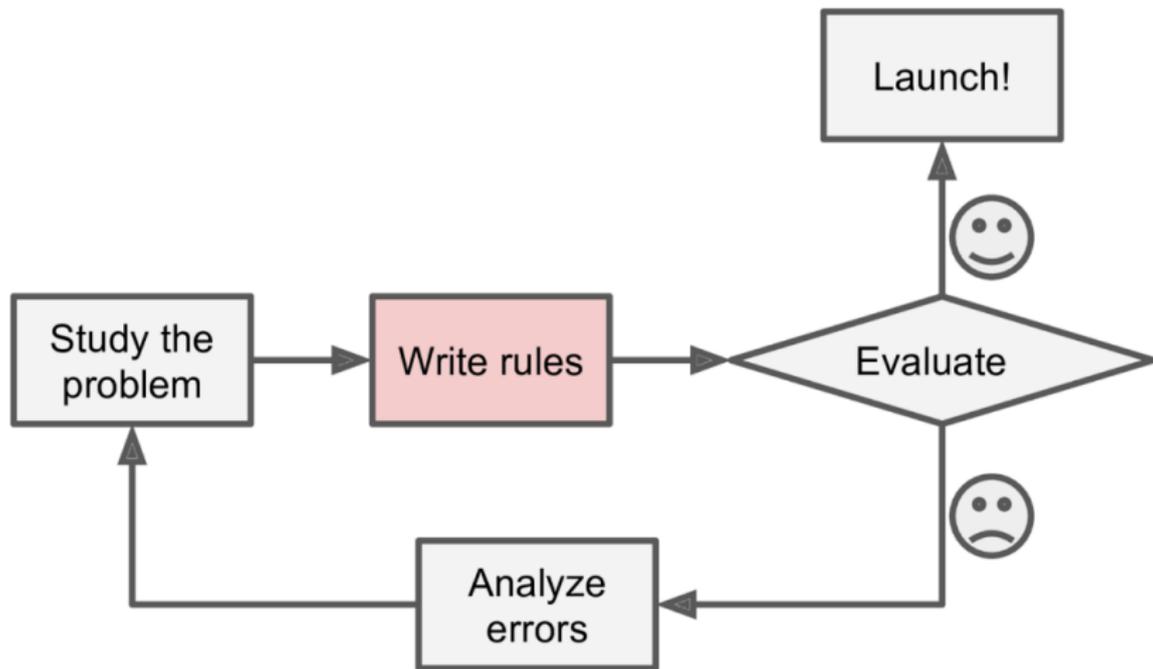


# Machine Learning

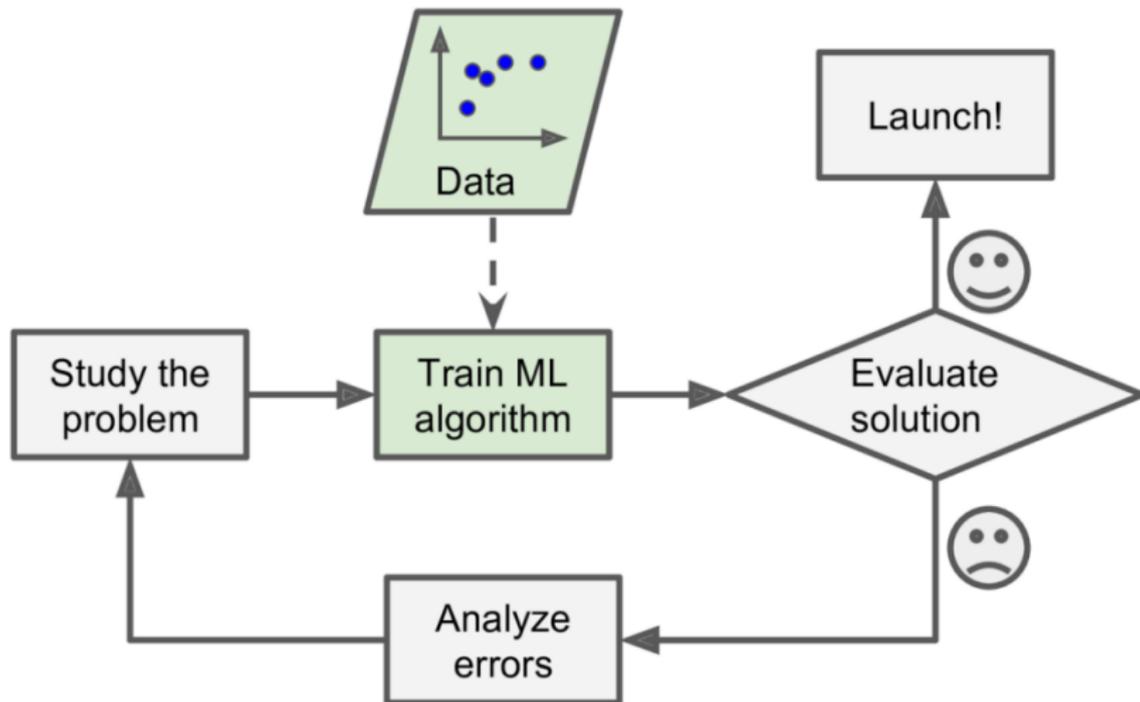
- A prediction function,  $f$ , takes some input,  $\mathbf{X}$ , and produces a meaningful output,  $\mathbf{Y}$
- Goal of machine learning is to *learn* the prediction function
  - Known inputs and outputs used to train the model
- Patterns in the data not obvious to us can be recognized in the learning process



# Traditional (Human-Based) Analysis Approach



# Machine Learning Analysis Approach



# Machine Learning

- Two broad categories of machine learning
  - Supervised learning
    - Provide training data with explicitly labelled inputs and outputs
  - Unsupervised learning
    - Provide unlabelled data and try to infer patterns from it
- Divide supervised learning into two categories based on the target(s)

## Classification (binary/multiclass)



0 0 0 0 0 0 0 0 0 0 0 0 0 0 0  
1 1 1 1 1 1 1 1 1 1 1 1 1 1 1  
2 2 2 2 2 2 2 2 2 2 2 2 2 2 2  
3 3 3 3 3 3 3 3 3 3 3 3 3 3 3  
4 4 4 4 4 4 4 4 4 4 4 4 4 4 4  
5 5 5 5 5 5 5 5 5 5 5 5 5 5 5  
6 6 6 6 6 6 6 6 6 6 6 6 6 6 6  
7 7 7 7 7 7 7 7 7 7 7 7 7 7 7  
8 8 8 8 8 8 8 8 8 8 8 8 8 8 8  
9 9 9 9 9 9 9 9 9 9 9 9 9 9 9

[3]

## Regression (continuous)



[4]

# Machine Learning for SNO+

- Machine learning can help in SNO+ data analysis
  - Antineutrino Search
    - *Classification* problem
    - Lots of data
    - Identify patterns / detect anomalies
  - Reconstruction of event positions in the detector
    - *Regression* problem
    - Lots of data
- Primary focus on neural networks
  - Found to perform faster and better than other approaches
  - Existing efficient implementations of matrix multiplication
  - Easy implementation of algorithms that utilize GPUs for training and inference

# Identifying Antineutrino Events

# AmBe Calibration Source

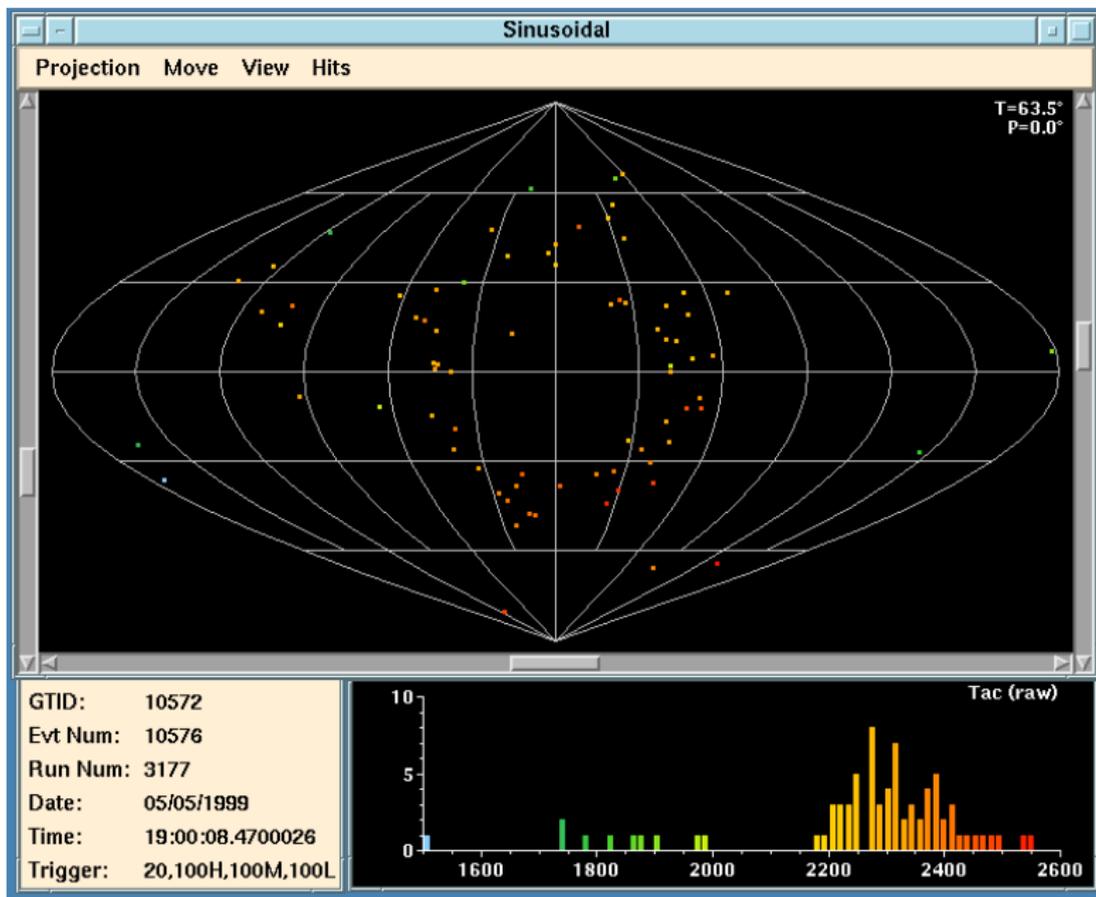
- Need a large training set of data for machine learning
- Americium-241 Beryllium-9, “AmBe”, calibration source
  - Provides two energy calibration measurements
  - **Mimics the antineutrino signal** ( $\bar{\nu}_e + p \rightarrow e^+ + n$ )



- 4.4MeV  $\gamma$  is basically instant (*prompt* event, like the  $e^+$ )
- n is captured in water some time later (*delayed* event)
  - Emits 2.2MeV  $\gamma$
  - Fitted neutron capture time in water:  
 $202.6 \pm 3.7\mu\text{s}$  [5],  $204.8 \pm 0.4\mu\text{s}$  [6],  $208.2 \pm 2.1\mu\text{s}$  [7]

# Neural Network Classifier: Approach

- Machine learning allows for the identification of the neutron capture signal
  - Supervised learning problem
  - Trained on AmBe Monte Carlo data
    - Salted with background data from the detector
  - Inputs: PMT positional coordinates and PMT time
    - PMT coordinates relative to the median of all hit PMTs in an event
    - Total size is  $4 \cdot nhits$  ( $x - x_m, y - y_m, z - z_m, t$ )
  - Output: a classification
    - Neutron capture or background

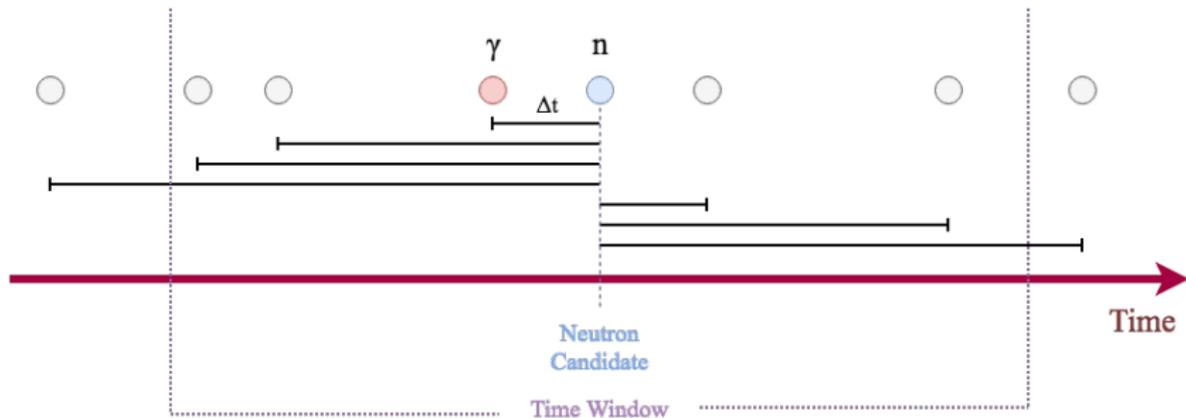


## Neural Network Classifier: Results

- Accuracy on test set  $>90\%$  for identification of both background and neutron capture
  - The test set is completely separated from the training set
- More importantly, predictions are consistent with real AmBe calibration data from the detector
  - Coincident signal is used to test the model
  - The time difference between delayed events predicted by the neural network (neutron candidates) and the event directly before (prompt event) should produce a specific timing distribution
  - Timing distribution fit will produce the neutron capture constant in water

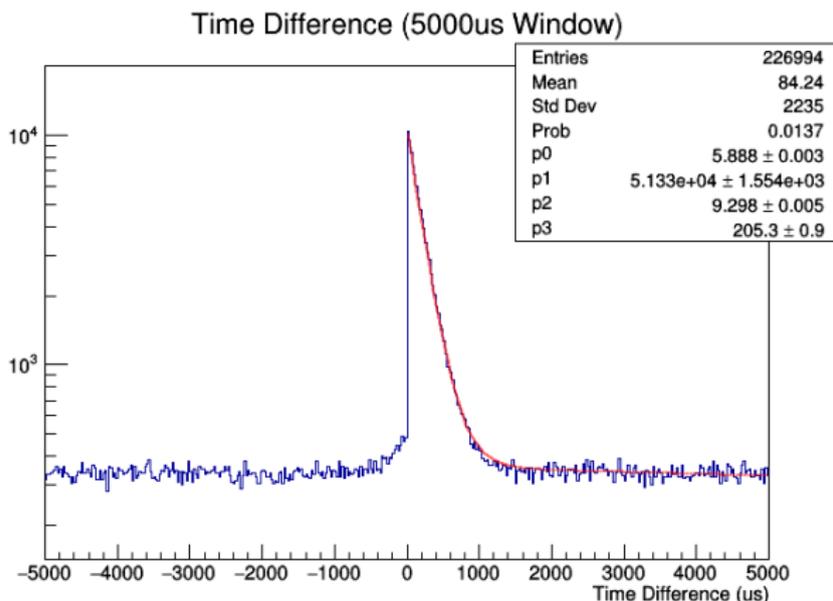
## Neural Network Classifier: Results

- Time difference between events within a  $5000\mu\text{s}$  before/after the candidate are binned
- Negative time difference (i.e., events *after* the candidate) are Poisson distributed similarly to data
  - Consists of backgrounds
- Positive time difference (i.e., events *before* the candidate) are Poisson distributed with a parameter of neutron capture constant
  - Consists of prompt events and backgrounds



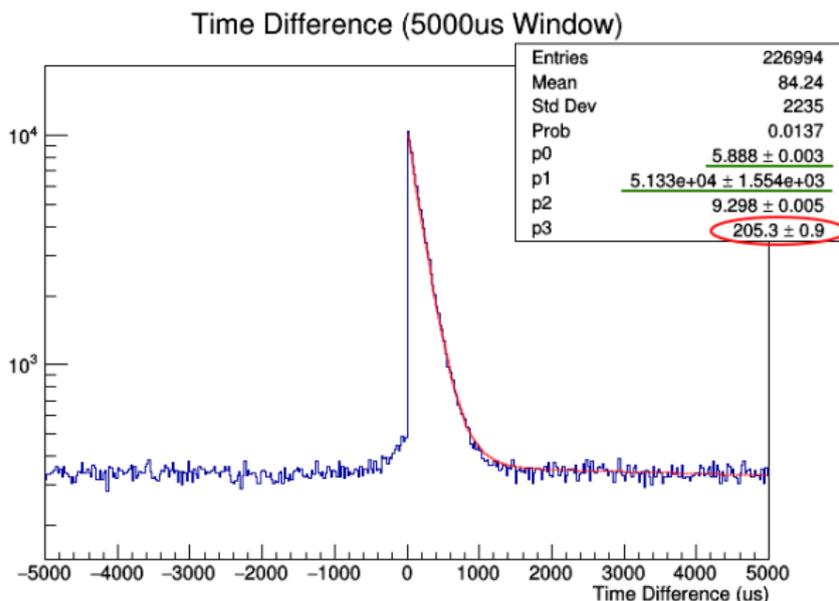
# Neural Network Classifier Performance

- Fit sum of two exponentials,  $\exp\left([0] - \frac{1}{[1]}\Delta t\right) + \exp\left([2] - \frac{1}{[3]}\Delta t\right)$ 
  - $p0$  and  $p1$  are constrained by negative fit and thus fixed
  - $p3$  is the neutron capture constant!



# Neural Network Classifier Performance

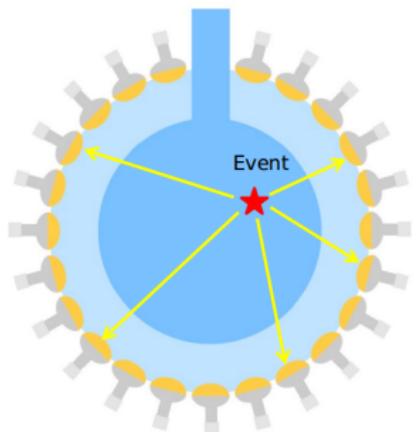
- Fit sum of two exponentials,  $\exp\left([0] - \frac{1}{[1]}\Delta t\right) + \exp\left([2] - \frac{1}{[3]}\Delta t\right)$ 
  - $p0$  and  $p1$  are constrained by negative fit and thus fixed
  - $p3$  is the neutron capture constant!



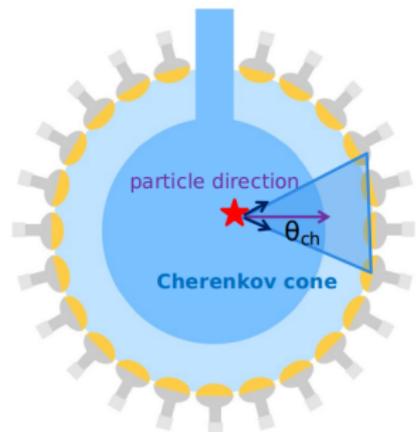
# Reconstruction of Event Position in the Detector

# Position Reconstruction

- Position reconstruction also suitable for machine learning
- Want to figure out where in the detector an event occurs
- Use relative PMT timing information
- Current approach involves a cascading of likelihood algorithms that account for complex physics processes



[8]

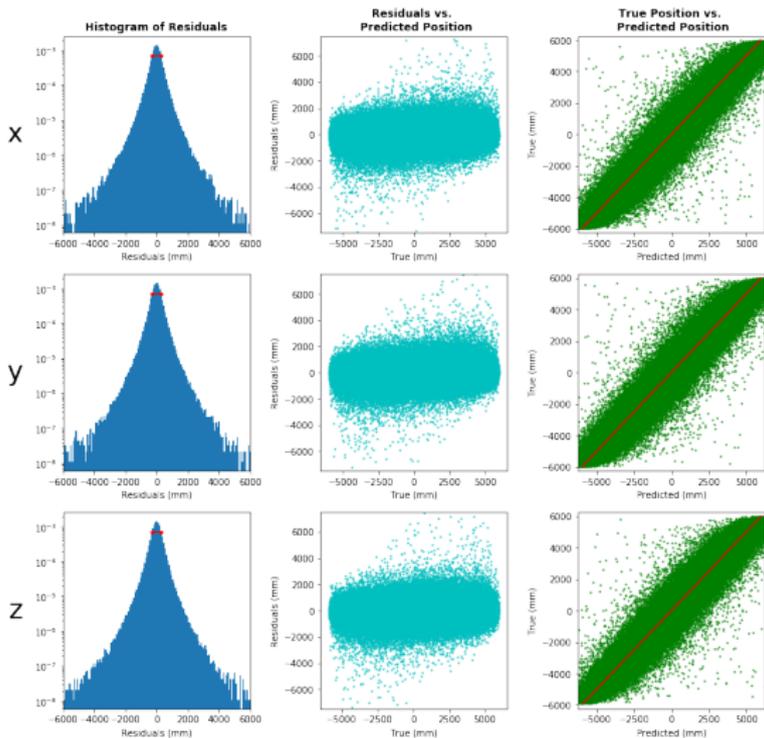


[8]

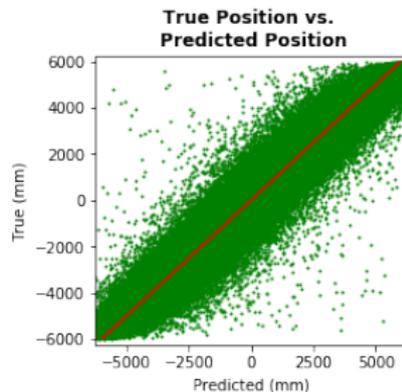
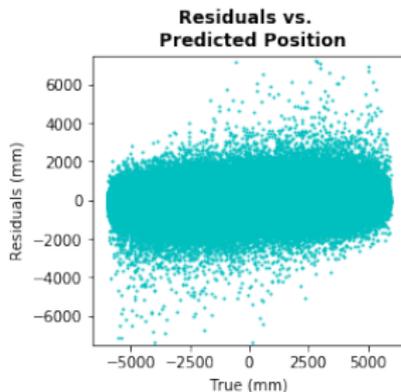
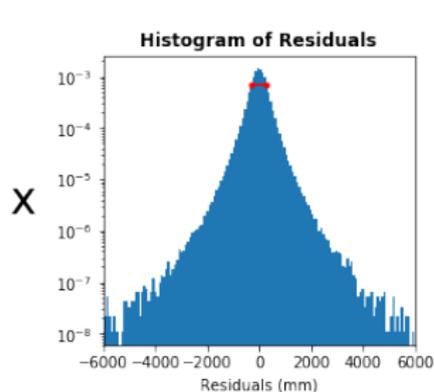
# Position Reconstruction

- With neural networks, a computational investment initially allows for fast inference later
  - Existing efficient implementations of matrix multiplication
  - Easy implementation of algorithms that take advantage of GPUs
  - A trained network is much quicker
- Supervised learning, regression problem
  - Train on Monte Carlo electrons
    - Uniformly distributed throughout the detector volume
  - Test on Monte Carlo and calibration data (tagged sources)

# Position Reconstruction: SNO+



# Position Reconstruction: SNO+

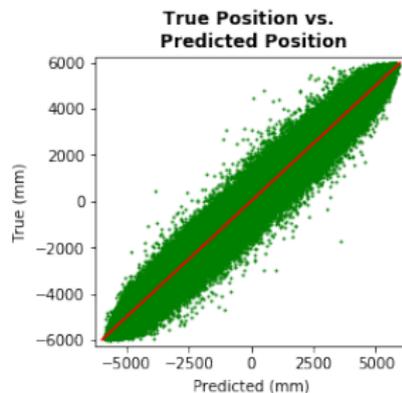
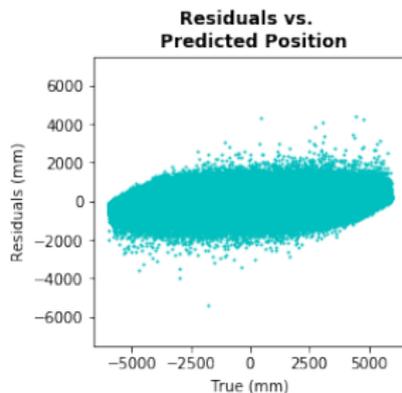
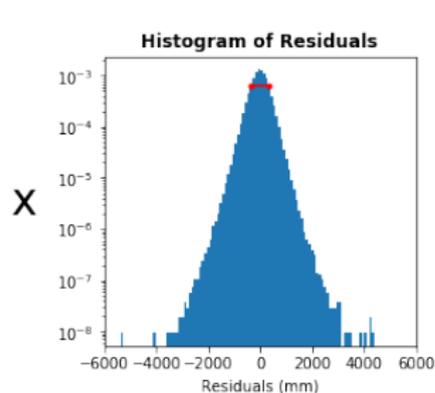


$$\mu_{\text{data}} = -0.261$$

$$\sigma_{\text{data}} = 399.312$$

$$\text{FWHM} = 586.344$$

# Position Reconstruction: Neural Network

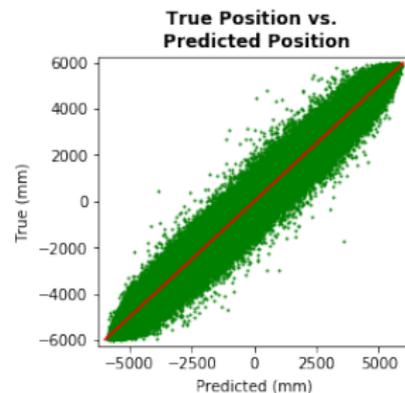
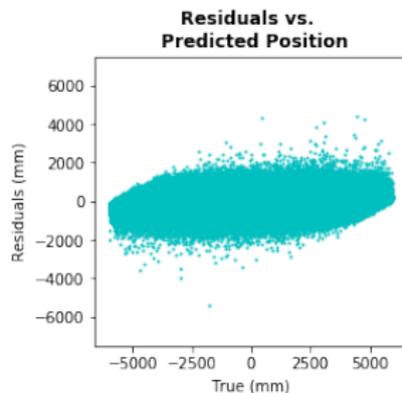
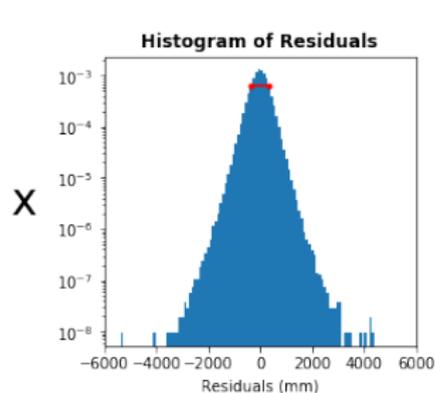


$$\mu_{\text{data}} = -9.796$$

$$\sigma_{\text{data}} = 350.687$$

$$\text{FWHM} = 675.440$$

# Position Reconstruction: Neural Network



$$\mu_{\text{data}} = -9.796$$

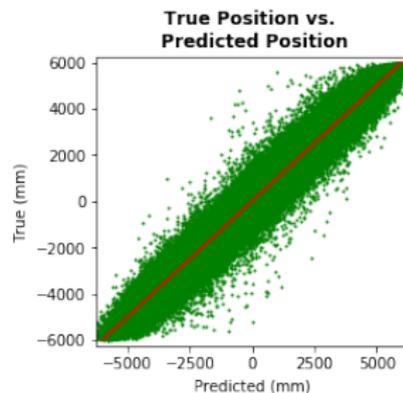
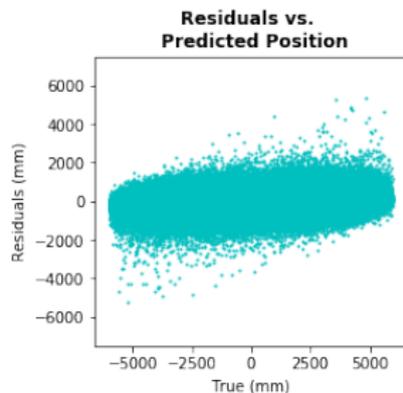
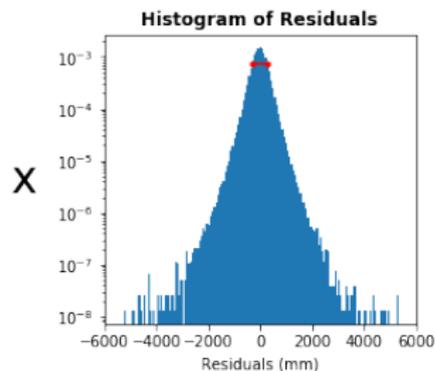
$$\sigma_{\text{data}} = 350.687$$

$$\text{FWHM} = 675.440$$

# Position Reconstruction

- Resolution of neural network is slightly worse
- Spread of residuals has less outliers
- Neural network makes much faster predictions than SNO+ fitter
  - 100x - 1,000x quicker on CPU (Central Processing Unit)
  - 10,000x quicker on GPU (Graphics Processing Unit)
  - Conservative estimates
- Average results together to obtain better reconstruction

# Position Reconstruction: Neural Network and SNO+

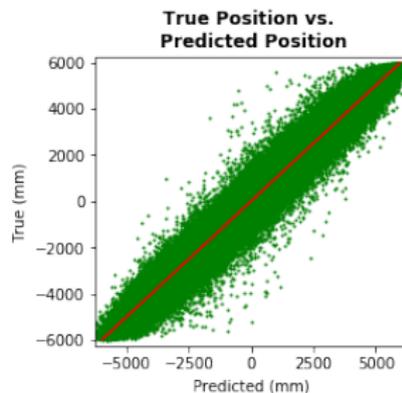
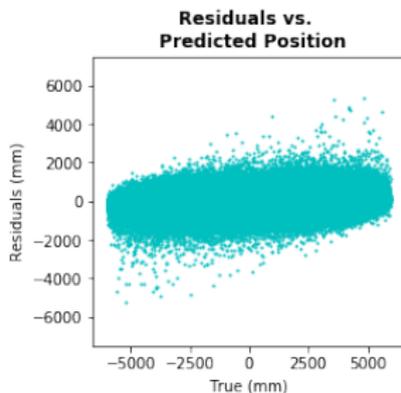
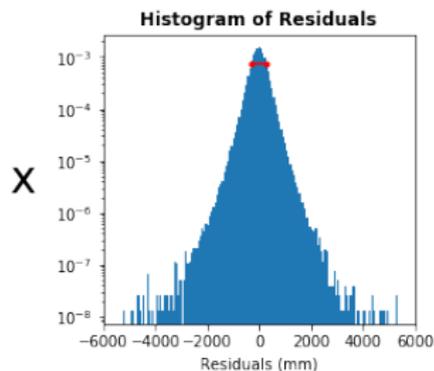


$$\mu_{\text{data}} = -5.028$$

$$\sigma_{\text{data}} = 326.874$$

$$\text{FWHM} = 580.510$$

# Position Reconstruction: Neural Network and SNO+



$$\mu_{\text{data}} = -5.028$$

$$\sigma_{\text{data}} = 326.874$$

$$\text{FWHM} = 580.510$$

## Conclusion

- Machine learning has proven successful on several analysis tasks
- Many more places where machine learning can be helpful
  - Monte Carlo simulations
  - Data cleaning and further background identification/reduction
- Results are promising entering the scintillator phase of the experiment

Thank You!

# References

- [1] "Beta Decay." [upload.wikimedia.org/wikipedia/commons/a/aa/Beta-minus\\_Decay.svg](https://upload.wikimedia.org/wikipedia/commons/a/aa/Beta-minus_Decay.svg).  
Accessed: 2019-03-11.
- [2] A. Géron, *Hands-on machine learning with Scikit-Learn and TensorFlow: concepts, tools, and techniques to build intelligent systems*.  
O'Reilly Media, Inc., 2017.
- [3] "Sample of MNIST Dataset." [commons.wikimedia.org/wiki/File:MnistExamples.png](https://commons.wikimedia.org/wiki/File:MnistExamples.png).  
Accessed: 2019-03-11.
- [4] "Chart Line." [pixabay.com/chart-line-line-chart-diagram-trend-148256/](https://pixabay.com/chart-line-line-chart-diagram-trend-148256/).  
Accessed: 2019-03-11.
- [5] Y. Zhang, K. Abe, Y. Haga, Y. Hayato, M. Ikeda, K. Iyogi, J. Kameda, Y. Kishimoto, M. Miura, S. Moriyama, *et al.*, "First measurement of radioactive isotope production through cosmic-ray muon spallation in super-kamiokande iv," *Physical Review D*, vol. 93, no. 1, p. 012004, 2016.
- [6] D. Cokinos and E. Melkonian, "Measurement of the 2200 m/sec neutron-proton capture cross section," *Physical Review C*, vol. 15, no. 5, p. 1636, 1977.
- [7] Y. Liu, S. Andringa, D. Auty, F. Barão, R. Bayes, E. Caden, C. Grant, J. Grove, B. Krar, A. LaTorre, *et al.*, "Neutron detection in the sno+ water phase," *arXiv preprint arXiv:1808.07020*, 2018.
- [8] J. Hu.  
From WNPPC 2017 Presentation.